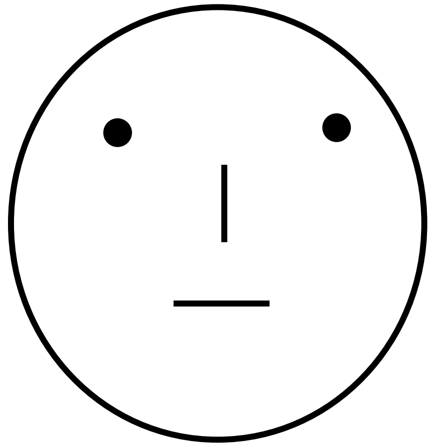
Computergrafik SS 2014
Oliver Vornberger

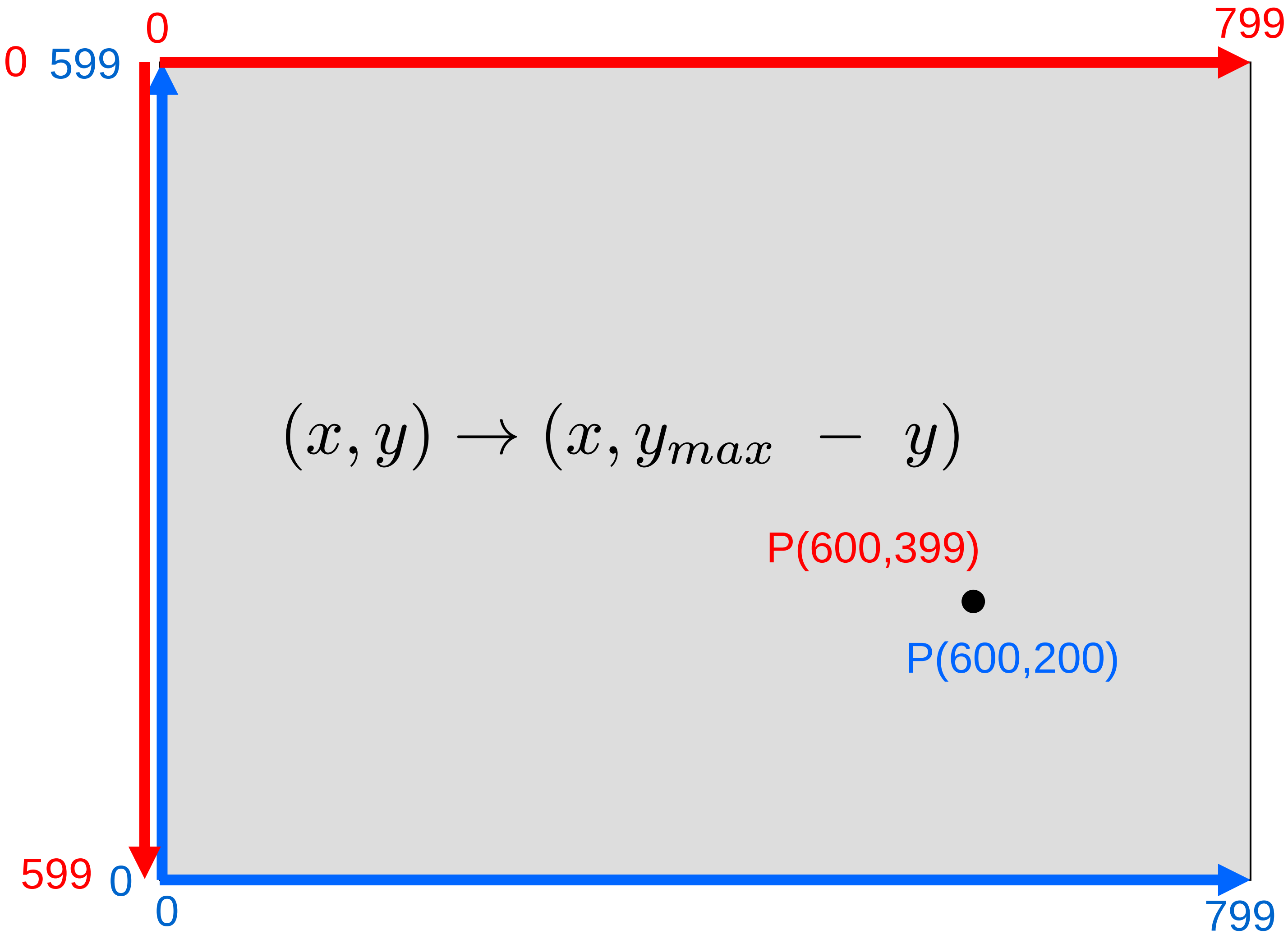
Kapitel 3:
2D-Grundlagen

Punkt, Punkt, Komma, Strich, ...

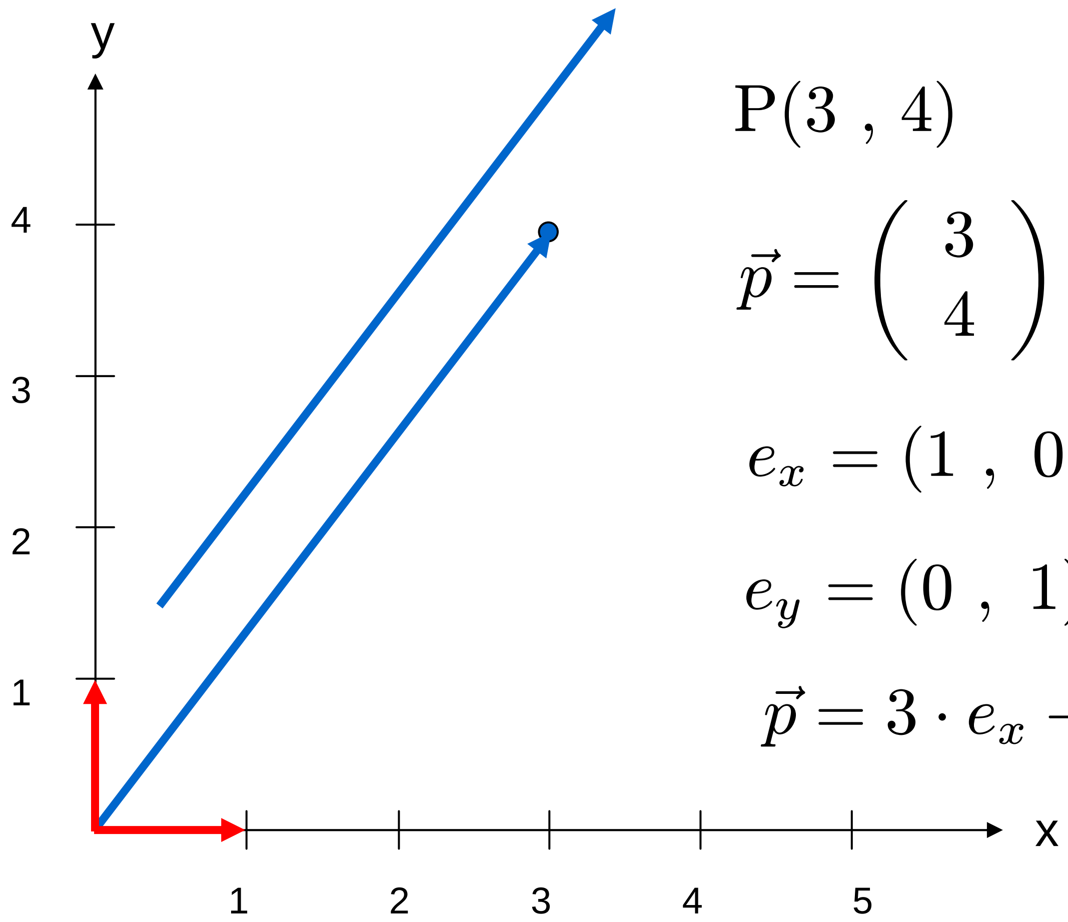


... fertig ist das Mondgesicht !

Koordinatensysteme



Punkt + Vektor



$$\mathrm{P}(3\ ,\ 4)$$

$$\vec{p} = \begin{pmatrix} 3 \\ 4 \end{pmatrix} = (3\ ,\ 4)^T$$

$$e_x = (1\ ,\ 0)^T$$

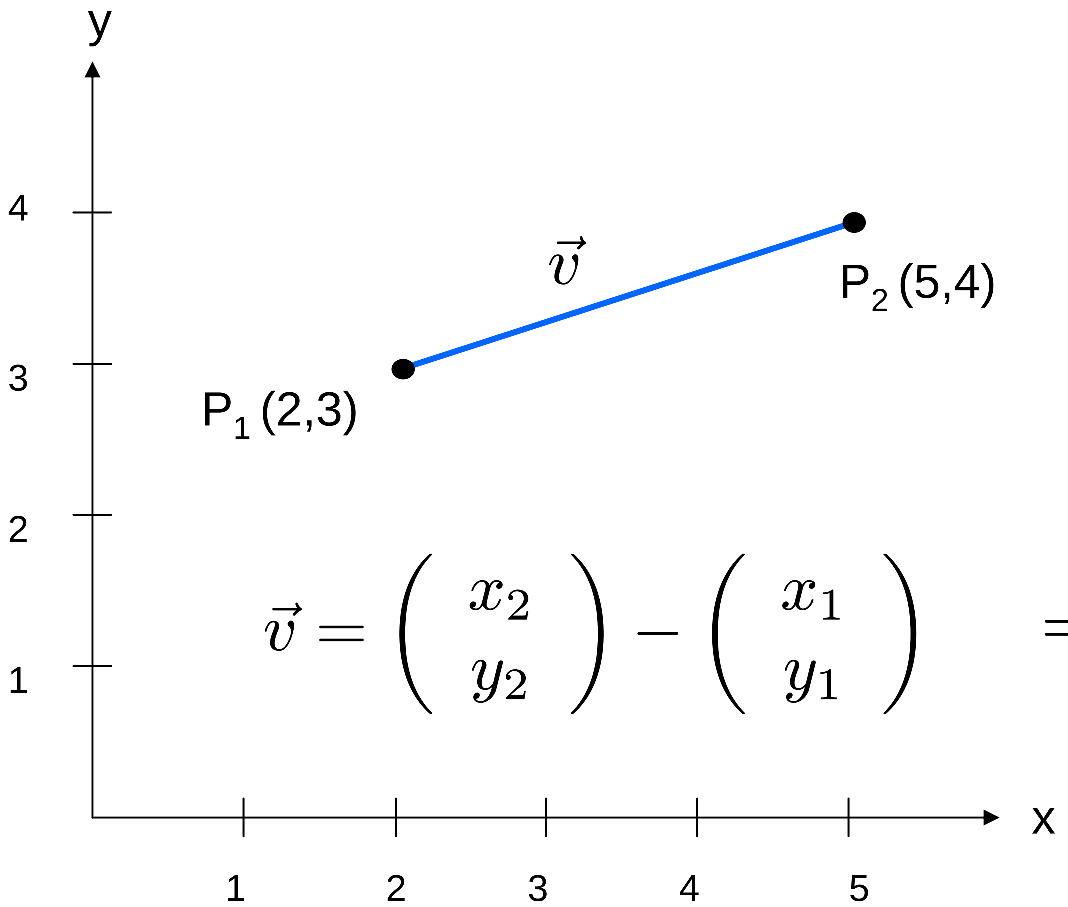
$$e_y = (0\ ,\ 1)^T$$

$$\vec{p} = 3 \cdot e_x + 4 \cdot e_y$$

setPixel(int x, int y)

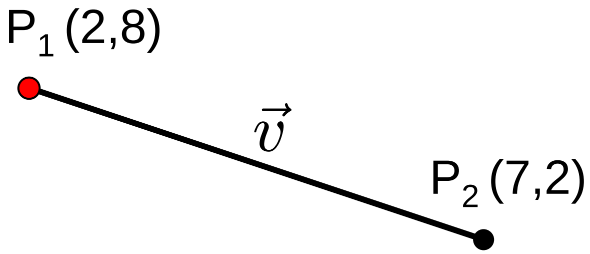
```
setPixel(3,4);

setPixel((int)(x+0.5),(int)(y+0.5));
```

Linie



$$\vec{v} = \begin{pmatrix} x_2 \\ y_2 \end{pmatrix} - \begin{pmatrix} x_1 \\ y_1 \end{pmatrix} = \begin{pmatrix} x_2 - x_1 \\ y_2 - y_1 \end{pmatrix}$$

Parametrisierte Gradengleichung



$$g : \vec{u} = \vec{p_1} + r \cdot \vec{v};\ \ r \in \mathbb{R}$$

$$l : \vec{u} = \vec{p_1} + r \cdot \vec{v};\ \ r \in [0;1]$$

1.0000

$$P = (1-t) \cdot P_1 + t \cdot P_2$$

$$d = \| \overline{P_1 P_2} \| \quad = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2}$$

$$step = \frac{1}{\sqrt{(x_2-x_1)^2+(y_2-y_1)^2}}$$

VectorLine

```
int x1,y1,x2,y2,x,y,dx,dy;
double r, step;

dy = y2-y1;
dx = x2-x1;

step = 1.0/Math.sqrt(dx*dx+dy*dy);
for (r=0.0; r <= 1; r=r+step) {
  x = (int)(x1+r*dx+0.5);
  y = (int)(y1+r*dy+0.5);
  setPixel(x,y);
}
```

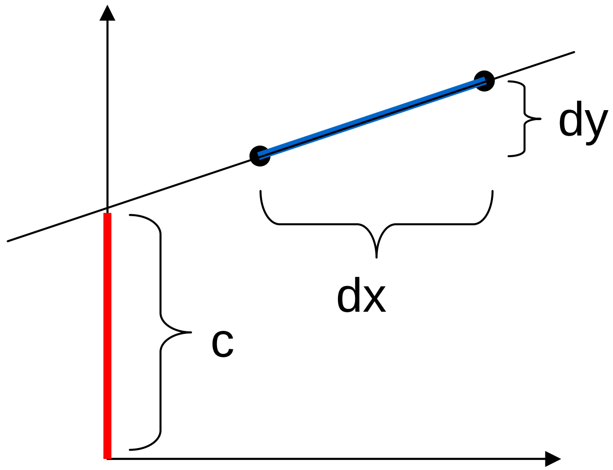
Gradengleichung als Funktion

$$y = f(x) = s \cdot x + c$$

$$s = \frac{y_2 - y_1}{x_2 - x_1}$$

$$\frac{y_1 - c}{x_1 - 0} = \frac{y_2 - y_1}{x_2 - x_1}$$

$$c = \frac{y_1 \cdot x_2 - y_2 \cdot x_1}{x_2 - x_1}$$



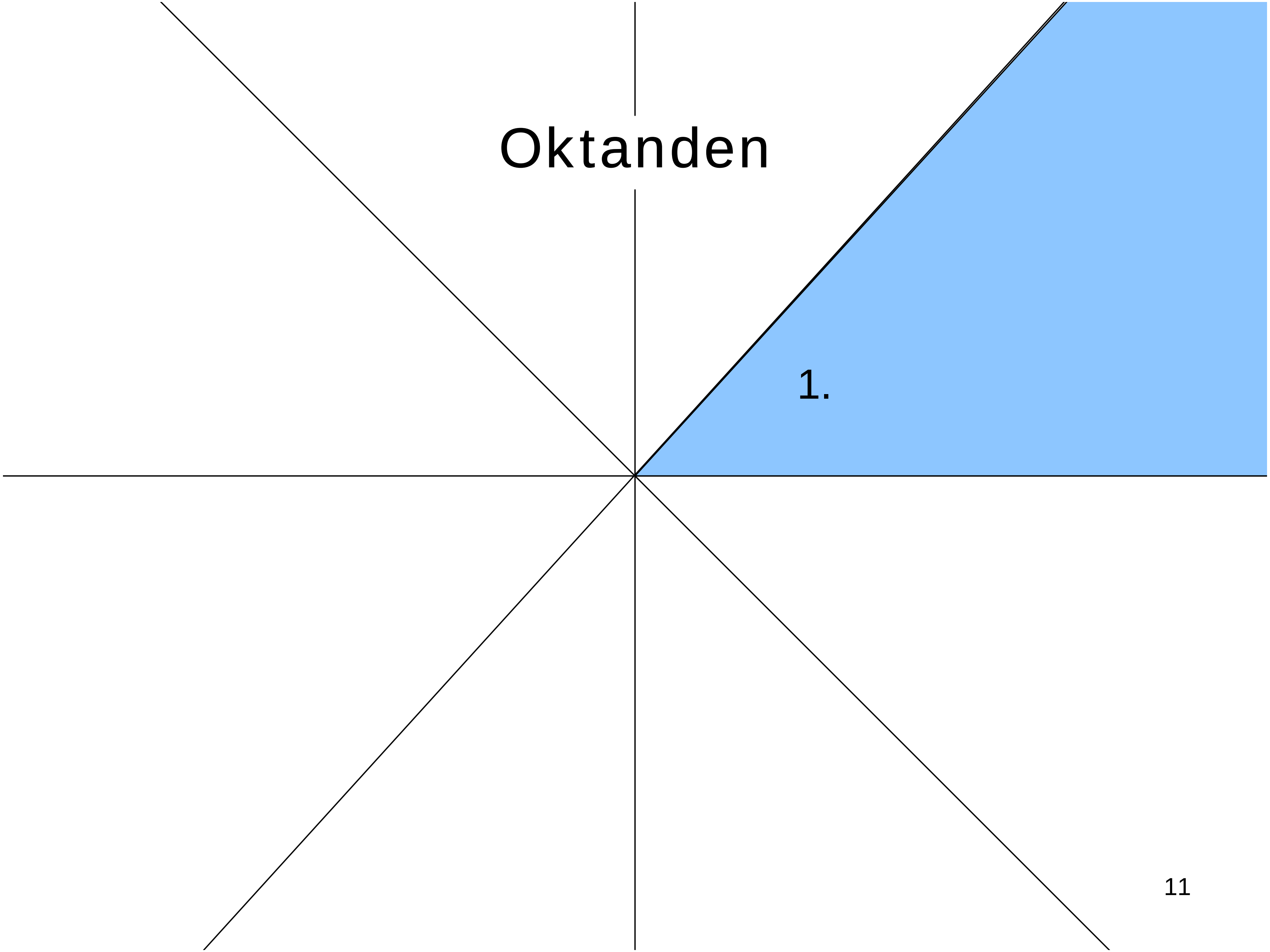
$$y = \frac{y_2 - y_1}{x_2 - x_1} \cdot x + \frac{x_2 \cdot y_1 - x_1 \cdot y_2}{x_2 - x_1}$$

StraightLine

von links nach rechts

```
s = (double)(y2-y1)/(double)(x2-x1);
c = (double)(x2*y1-x1*y2)/(double)(x2-x1);

for (x=x1; x <= x2; x++) {
  y = (int)(s*x+c+0.5);
  setPixel(x,y);
}
```

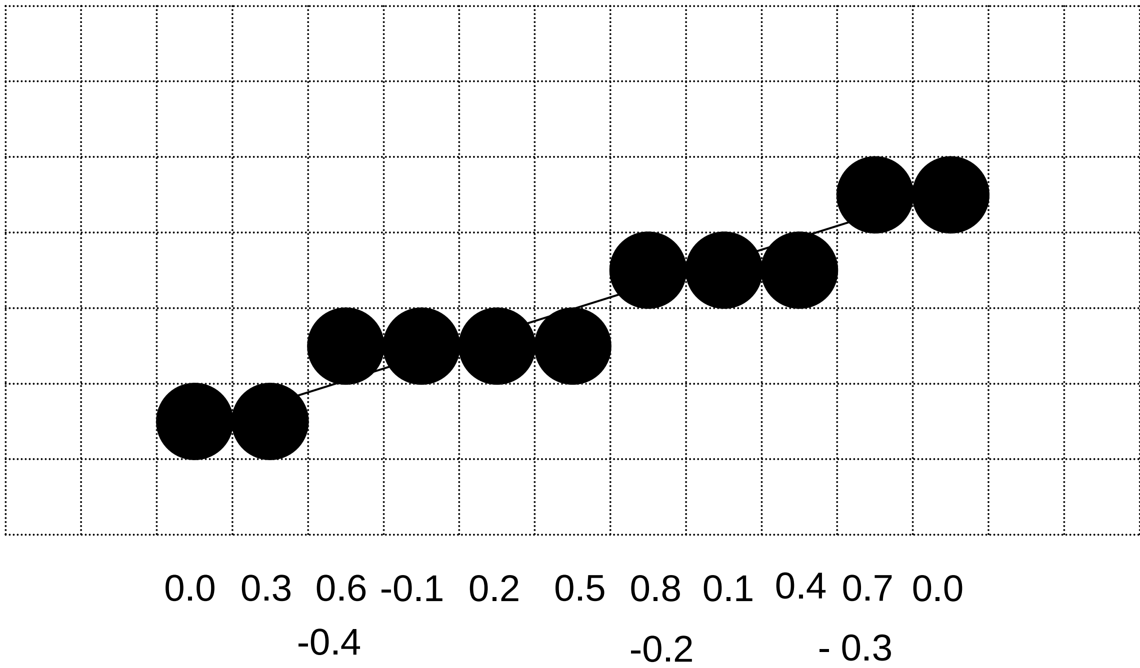
Oktanden



Bresenham

Steigung $s = \Delta y / \Delta x = 3/10 = 0.3$

Fehler $error = y_{ideal} - y_{real}$



BresenhamLine, die 1.

```
dy = y2-y1; dx = x2-x1;
s = (double)dy/(double)dx;
error = 0.0;
x = x1;
y = y1;
while (x <= x2){
  setPixel(x,y);
  x++;
  error = error + s;
  if (error > 0.5) {
    y++;
    error = error – 1.0;
  }
}
```

BresenhamLine

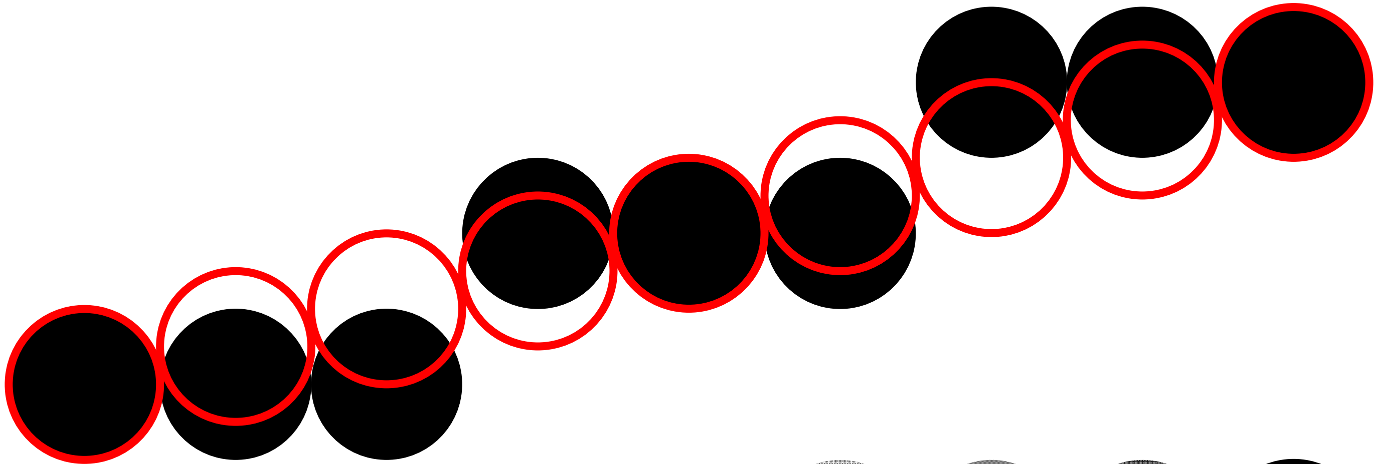
alle 8 Oktanden durch Fallunterscheidung abhandeln:

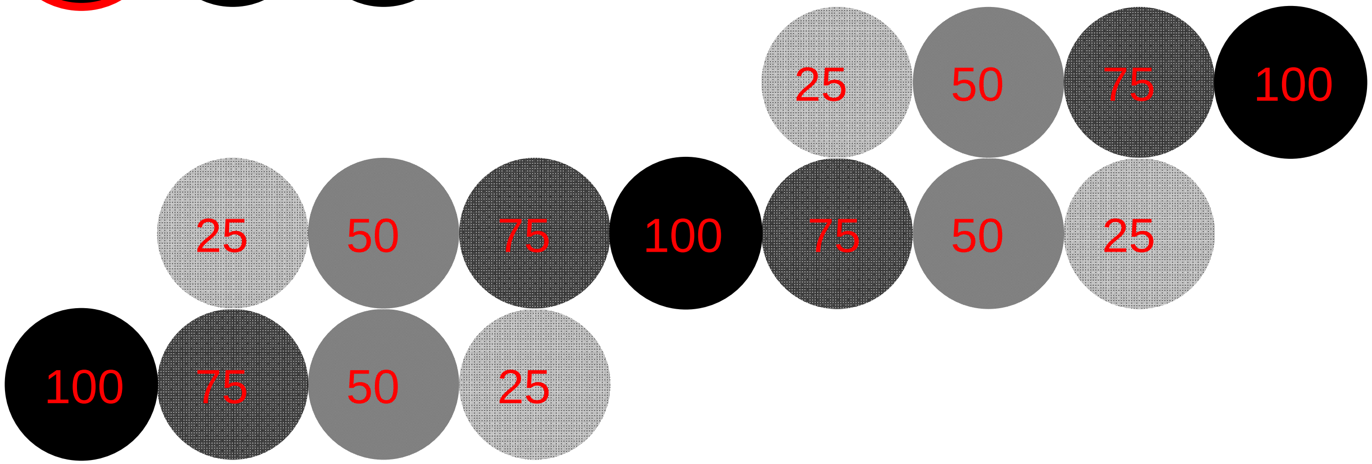
~cg/2014/skript/Sources/drawBresenhamLine.jav.html

Java-Applet:

~cg/2014/skript/Applets/2D-basic/App.html

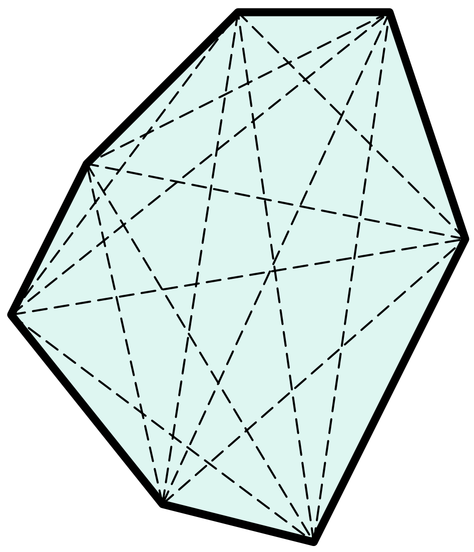
Antialiasing

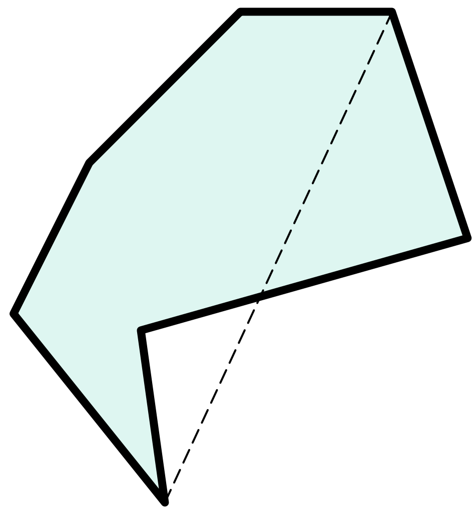




Antialiasing in Adobe Photoshop

Polygon





konvex konkav

Punkt versus Gerade

$$\begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} 2 \\ 3 \end{pmatrix} + r \cdot \begin{pmatrix} 7-2 \\ 5-3 \end{pmatrix} \qquad \vec{u} = \vec{p_1} + r \cdot \vec{v}$$

x = *2* + *5r*

y = *3* + *2r*

2x = *4* + *10r*

5y = *15* + *10r*

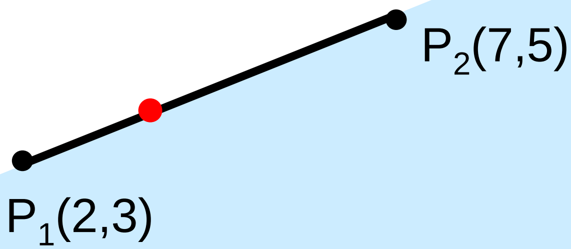
5y - *2x* = *11*

5y – *2x* - *11* = *0*

F(x,y) = *0* falls *P* auf der Geraden

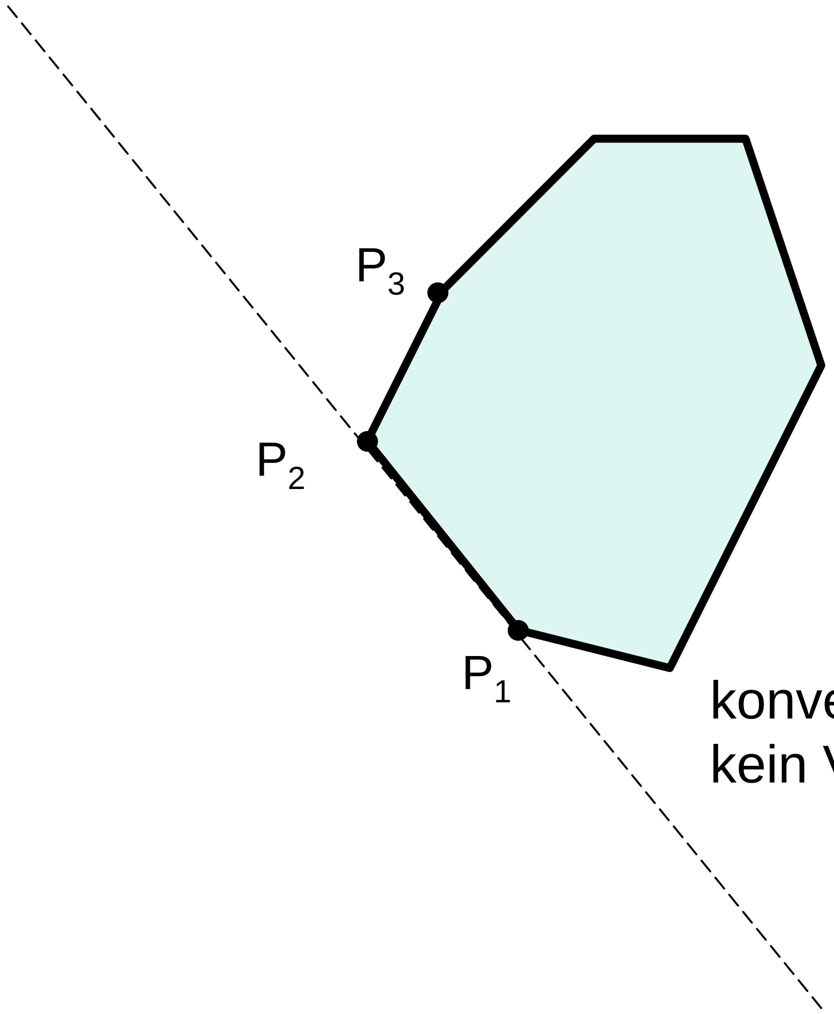
> *0* falls *P* links von der Geraden

< *0* falls *P* rechts von der Geraden



$$F(\vec{x}) = (\vec{x} - \vec{p_1}) \cdot \vec{n}$$

Konvexitätstest nach Paul Bourke



konvex, falls für alle P(x,y)
kein Vorzeichenwechsel bei F(x,y)

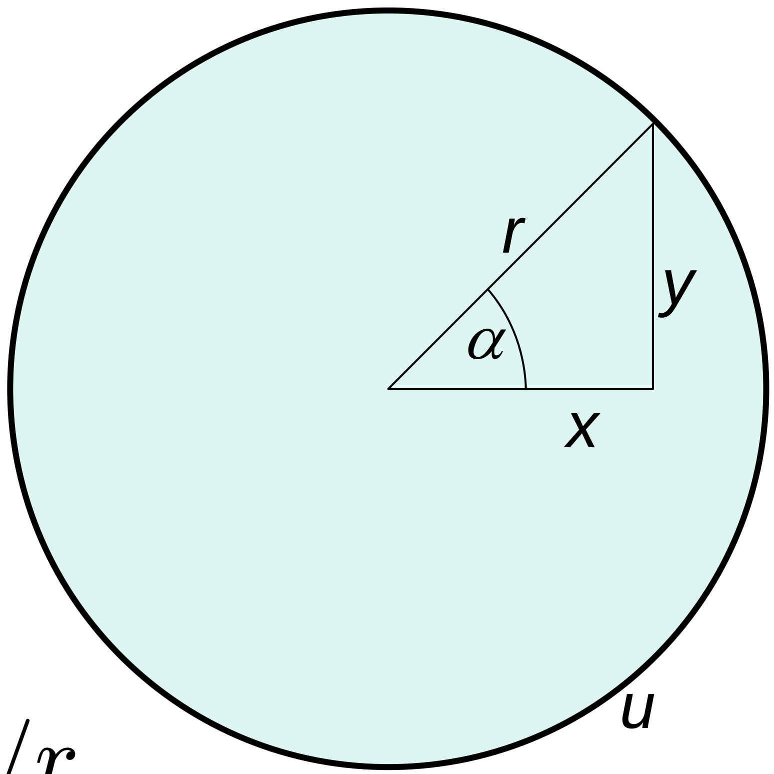
Kreis um (0,0), parametrisiert

$$x = r \cdot \cos(\alpha)$$

$$y = r \cdot \sin(\alpha)$$

$$u = 2 \cdot \pi \cdot r$$

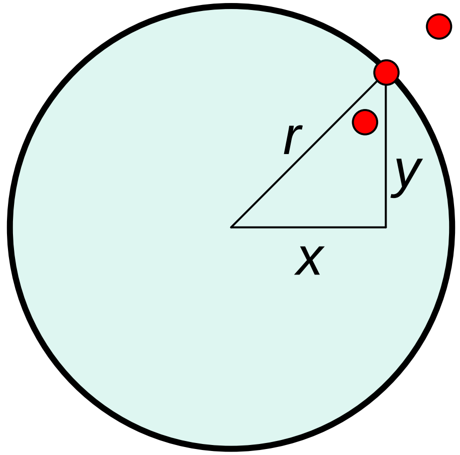
$$step = \frac{2 \cdot \pi}{2 \cdot \pi \cdot r} = 1/r$$



TriCalcCircle

```
double step = 1.0/(double r);
double winkel;

for (winkel = 0.0;
     winkel < 2*Math.PI;
     winkel = winkel+step){

  setPixel((int) r*Math.sin(winkel)+0.5,
           (int) r*Math.cos(winkel)+0.5);
}
```

Punkt versus Kreis



$$x^2 + y^2 = r^2$$

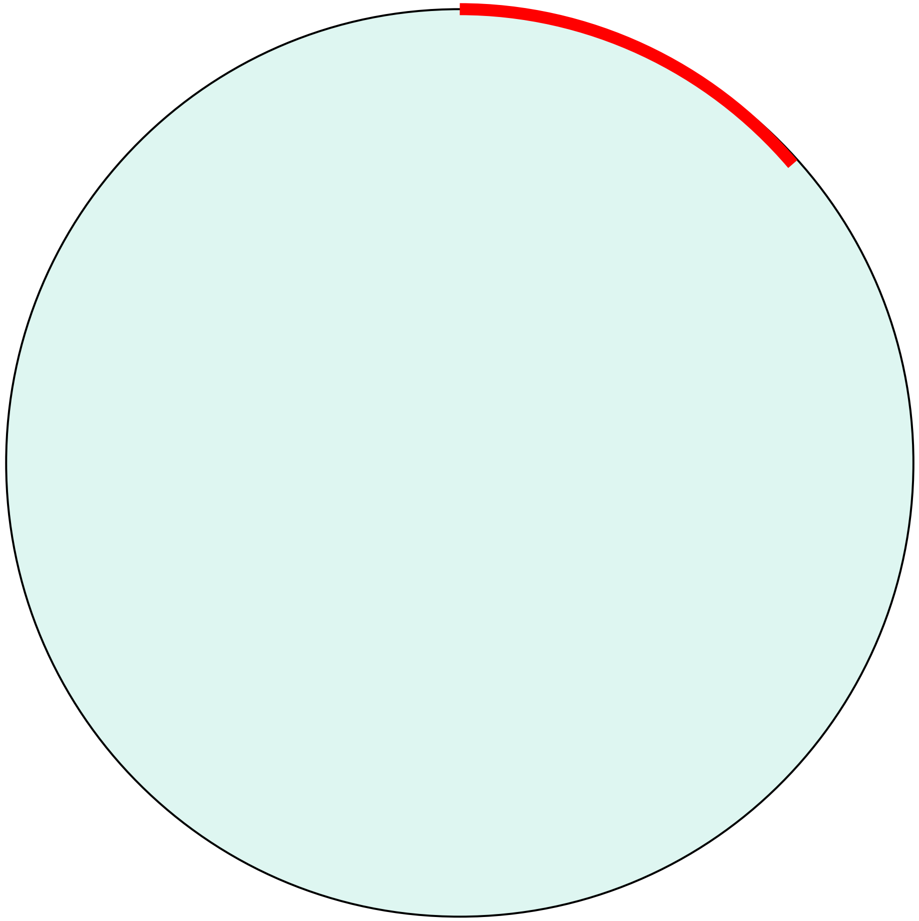
$$F(x,y) = x^2 + y^2 - r^2$$

$F(x,y)$ = 0 für (x,y) auf dem Kreis

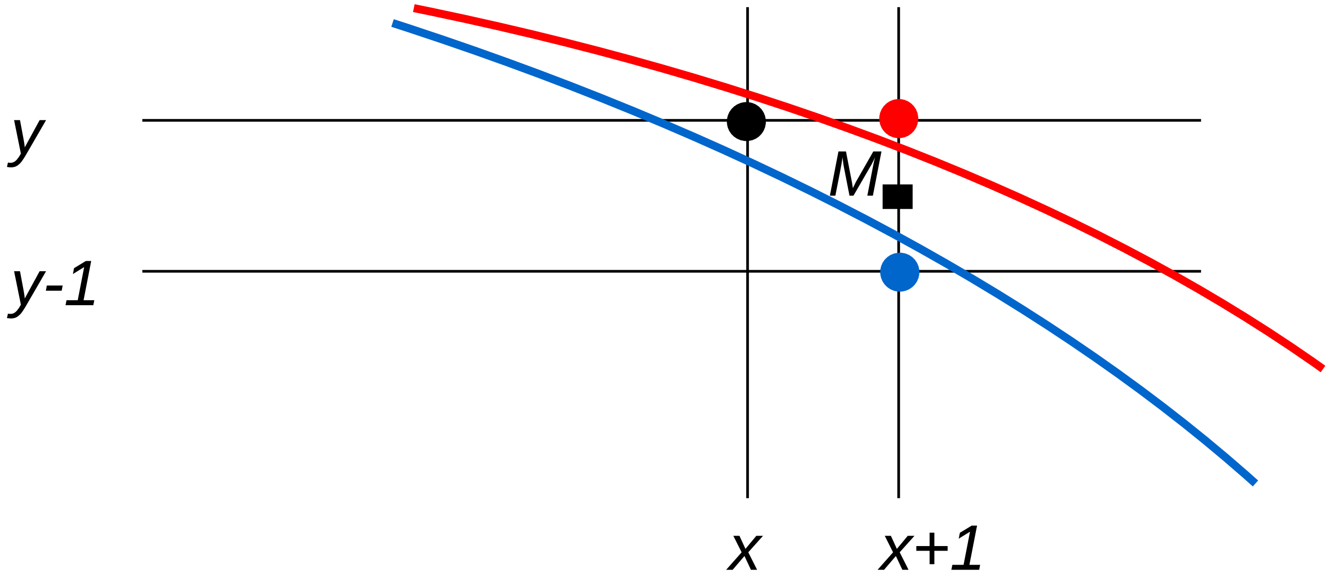
< 0 für (x,y) innerhalb des Kreises

> 0 für (x,y) außerhalb des Kreises

Kreis im 2. Oktanden



Entscheidungsvariable Δ



$\Delta = F(x+1, y-\frac{1}{2})$

$\Delta < 0 \Rightarrow$ *M* liegt innerhalb $\Rightarrow$ wähle $(x+1, y)$

$\Delta \geq 0 \Rightarrow$ *M* liegt außerhalb $\Rightarrow$ wähle $(x+1, y-1)$

Berechnung von Δ

$\Delta = F(x+1, y-½) = (x+1)^2+(y-½)^2-r^2$

$\Delta < 0 \Rightarrow$

$\Delta' = F(x+2, y-½) = (x+2)^2+(y-½)^2-r^2 =$

$\Delta+2x+3$

$\Delta \geq 0 \Rightarrow$

$\Delta' = F(x+2, y-3/2) = (x+2)^2+(y-3/2)^2-r^2 =$

$\Delta+2x-2y+5$

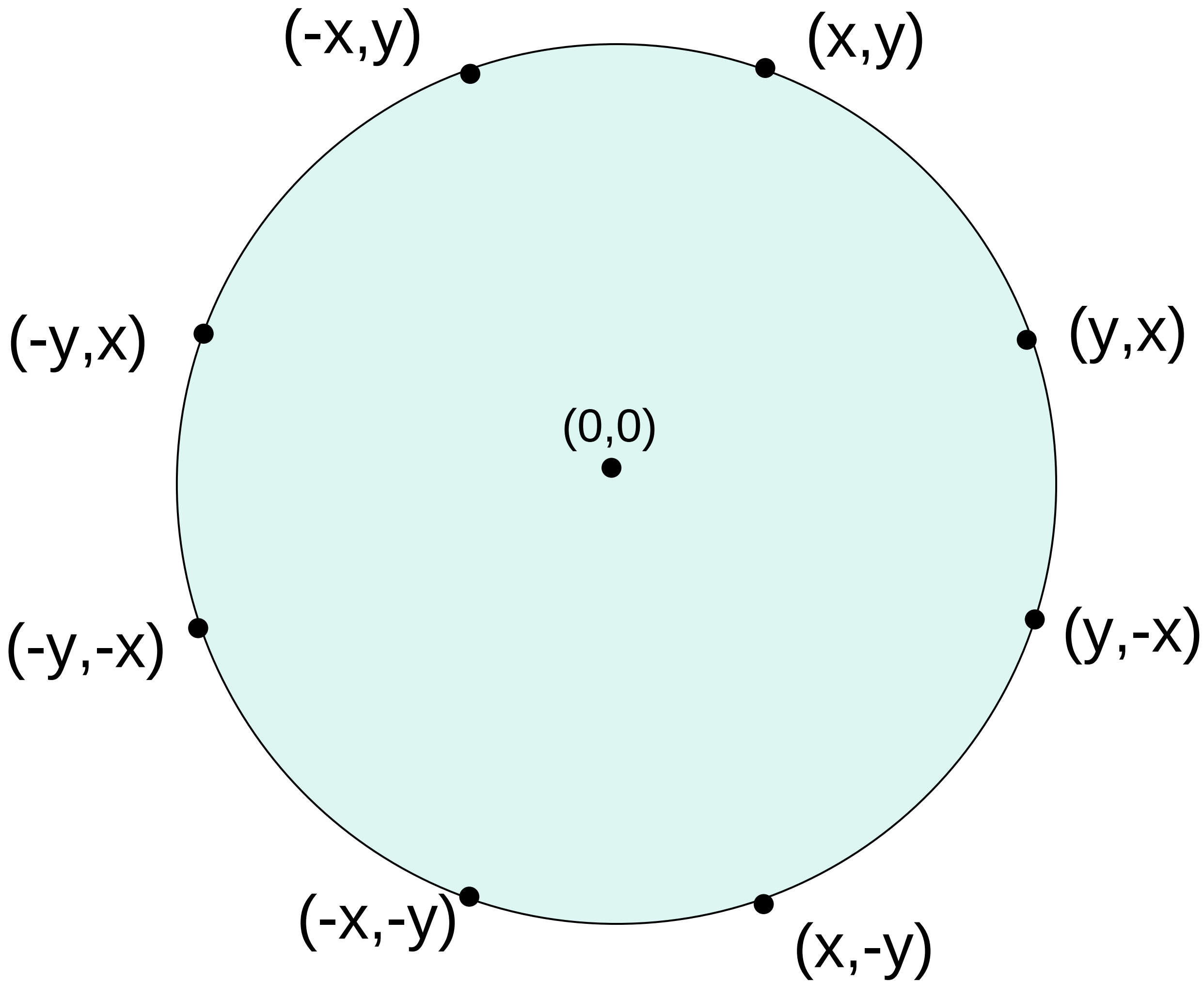
Startwert $\Delta = F(1, r-½) = 1^2+(r-½)^2 - r^2 =$

$5/4-r$

BresenhamCircle, die 1.

```
x = 0;
y = r;
delta = 5.0/4.0 - r;
while (y >= x) {
  setPixel(x,y);
  if (delta < 0.0) {
    delta = delta + 2*x + 3.0;
    x++;
  else {
    delta = delta + 2*x - 2*y + 5.0;
    x++;
    y--;
  }
}
```

Oktanden-Symmetrie



BresenhamCircle, die 3.

```
x=0; y=r; d=1-r; x=3; dx=3; dxy=-2*r+5;
while (y>=x){
   setPixel(+x,+y);
   setPixel(+y,+x);
   setPixel(+y,-x);
   setPixel(+x,-y);
   setPixel(-x,-y);
   setPixel(-y,-x);
   setPixel(-y,+x);
   setPixel(-x,+y);
  if (d<0) {d=d+dx;  dx=dx+2; dxy=dxy+2; x++;}
  else     {d=d+dxy; dx=dx+2; dxy=dxy+4; x++;
   y--;}
}
```

Source: [~cg/2014/skript/Sources/drawBresenhamCircle.jav]

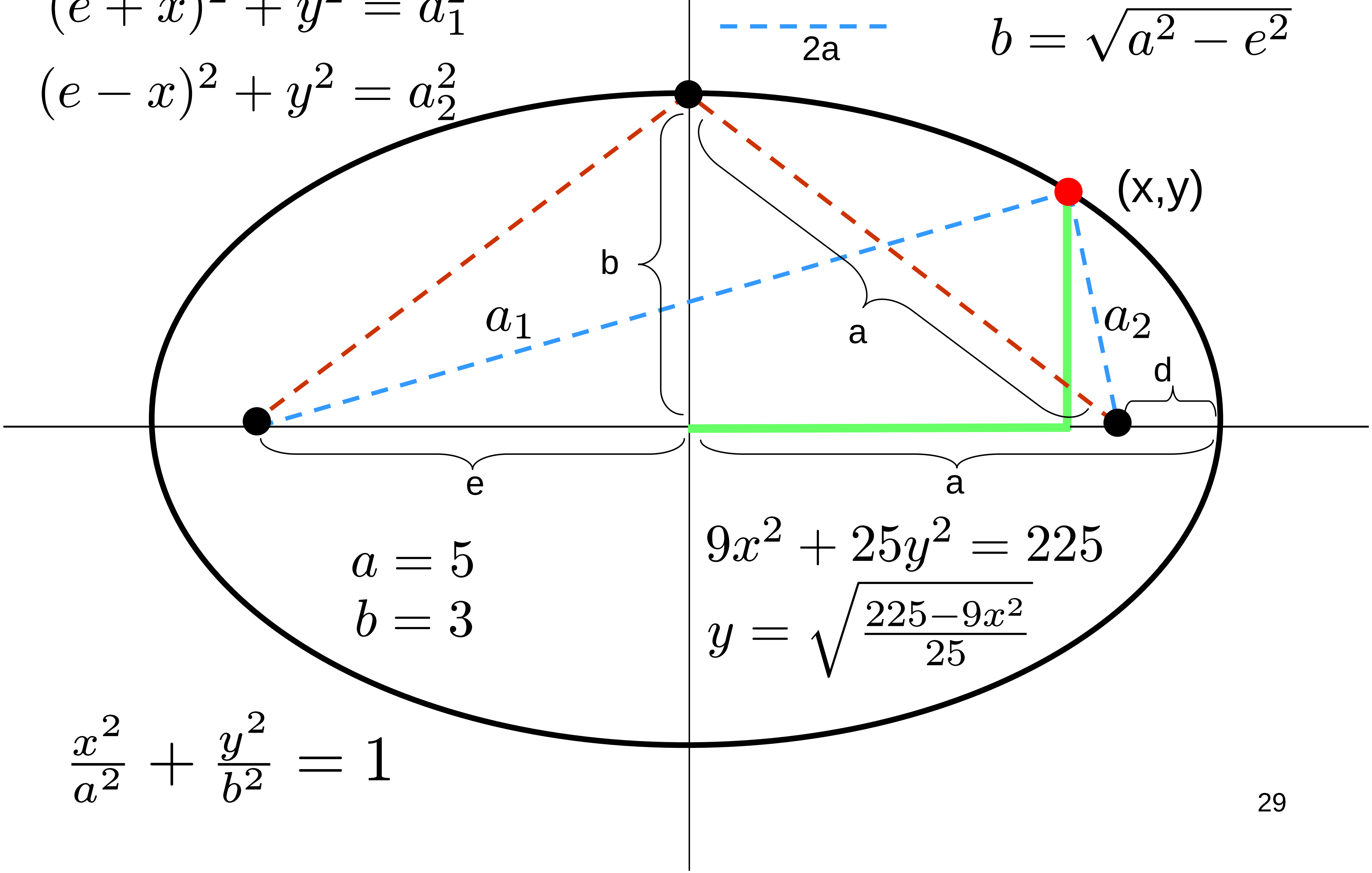
Java-Applet: [~cg/2014/skript/Applets/2D-basic/App.html]

Ellipse um (0,0)

$$(e+x)^2 + y^2 = a_1^2$$

$$(e-x)^2 + y^2 = a_2^2$$

$$b = \sqrt{a^2 - e^2}$$



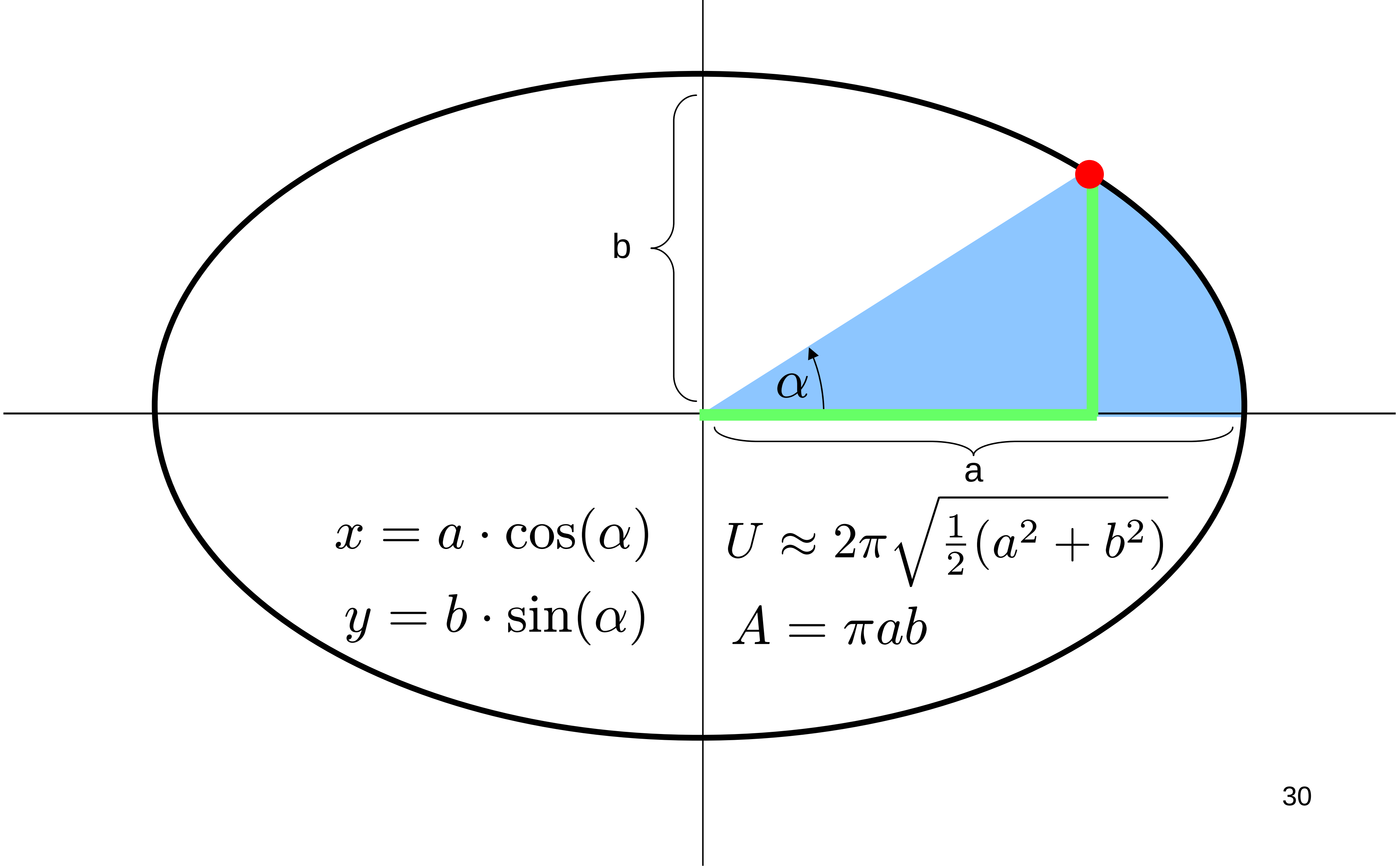
$$a = 5$$
$$b = 3$$

$$9x^2 + 25y^2 = 225$$

$$y = \sqrt{\frac{225-9x^2}{25}}$$

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$

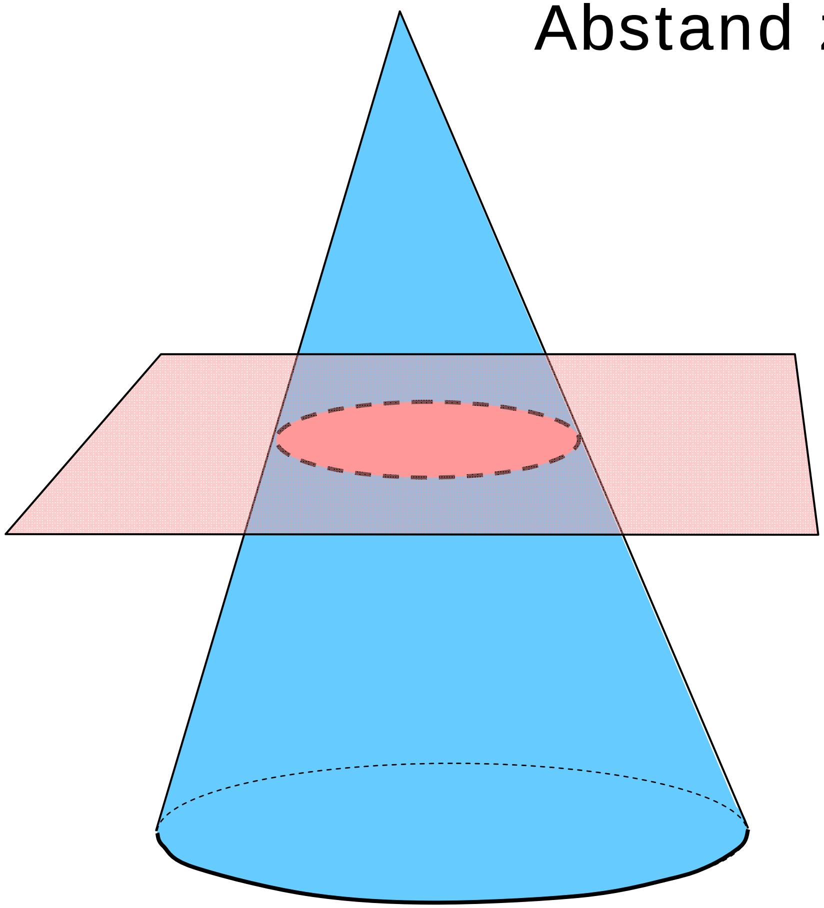
Ellipse um (0,0) parametrisiert



$$x = a \cdot \cos(\alpha)$$
$$y = b \cdot \sin(\alpha)$$

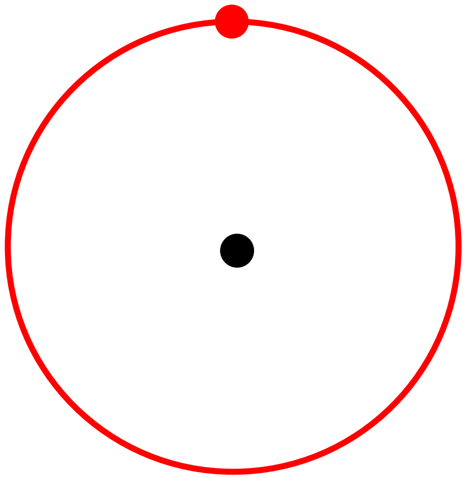
$$U \approx 2\pi \sqrt{\frac{1}{2}(a^2 + b^2)}$$
$$A = \pi a b$$

Kegelschnitt: Kreis

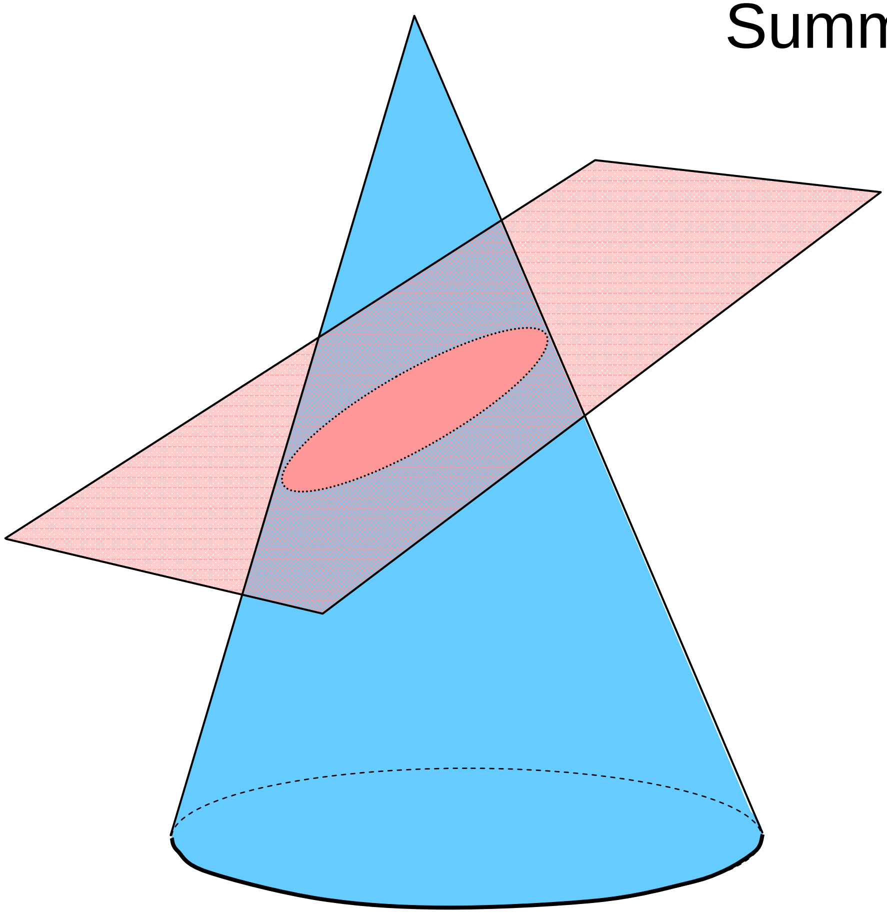


Abstand zu einem Punkt ist konstant

$$x^2 + y^2 = 1$$

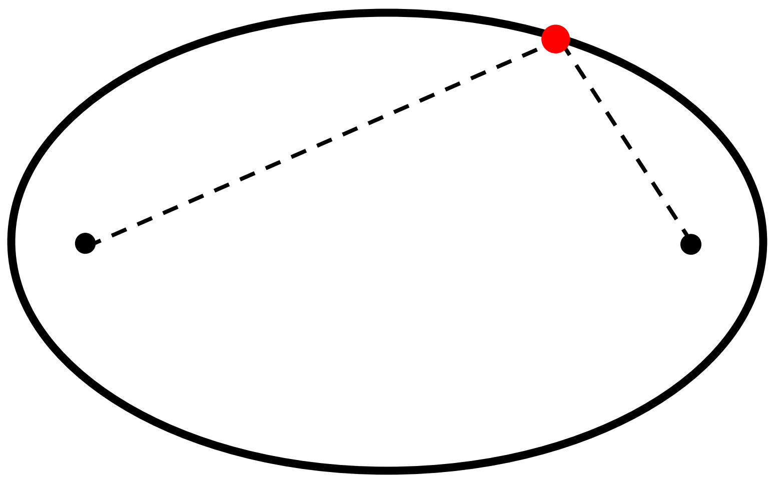


Kegelschnitt: Ellipse

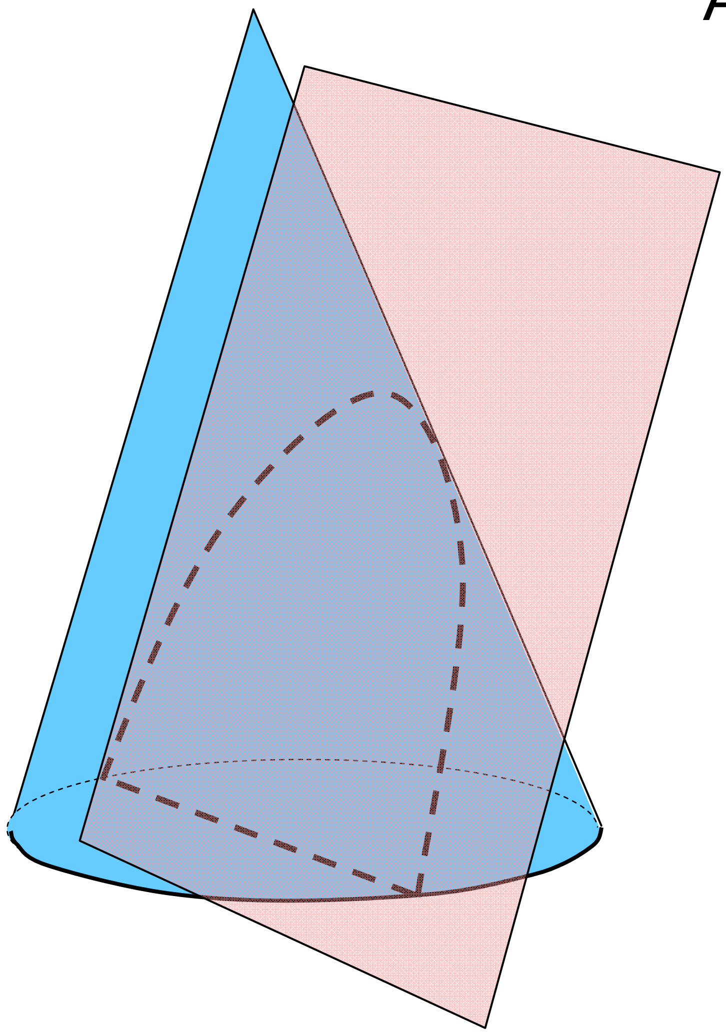


Summe der Abstände zu 2 Punkten
ist konstant

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$

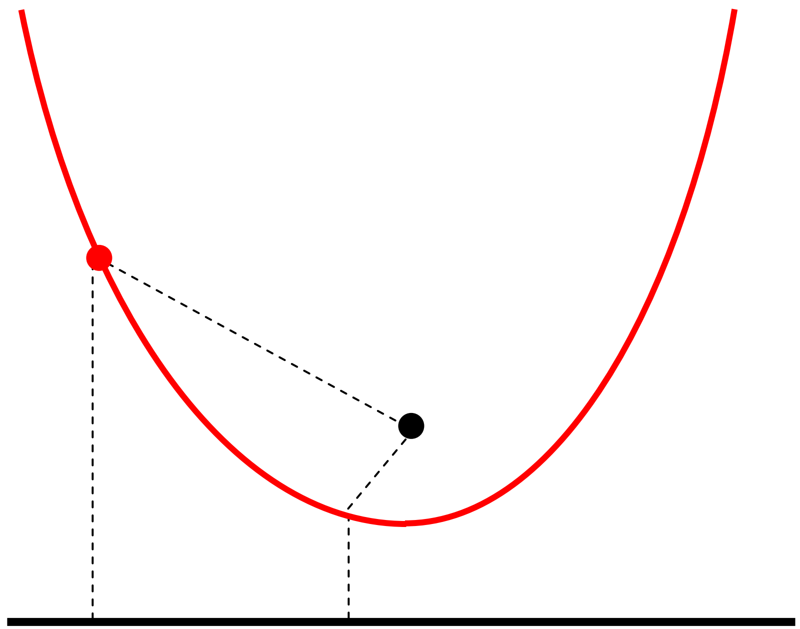


Kegelschnitt: Parabel



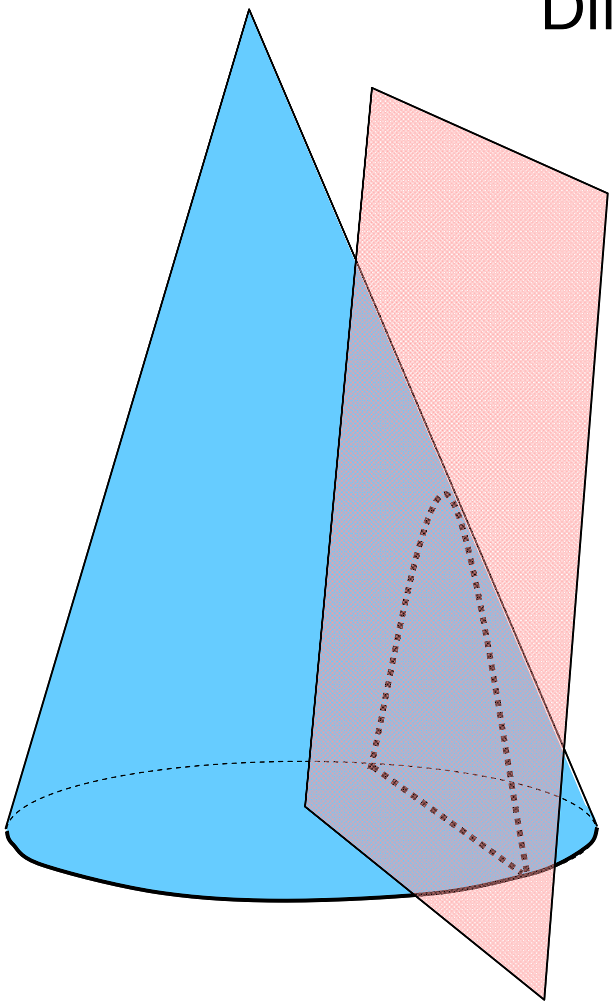
Abstand zu Punkt und Gerade
ist gleich

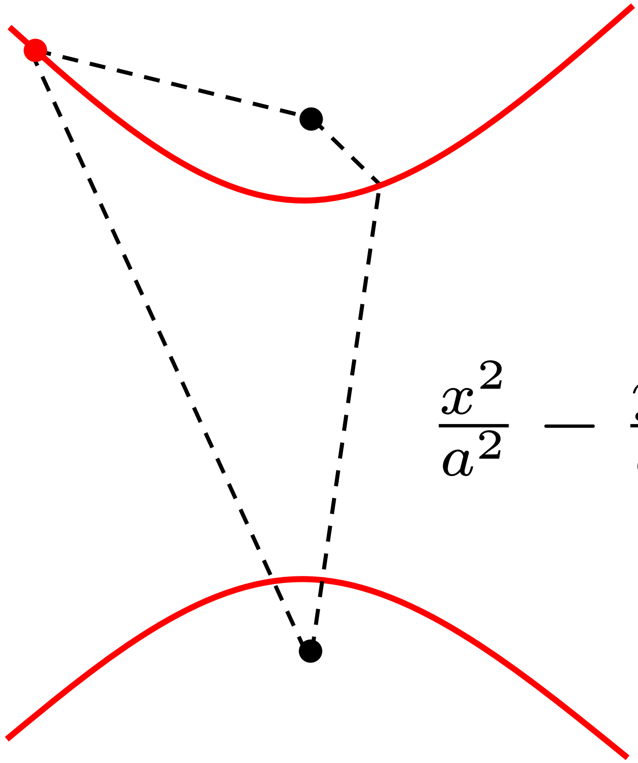
$$y = ax^2 + bx + c$$



Kegelschnitt: Hyperbel

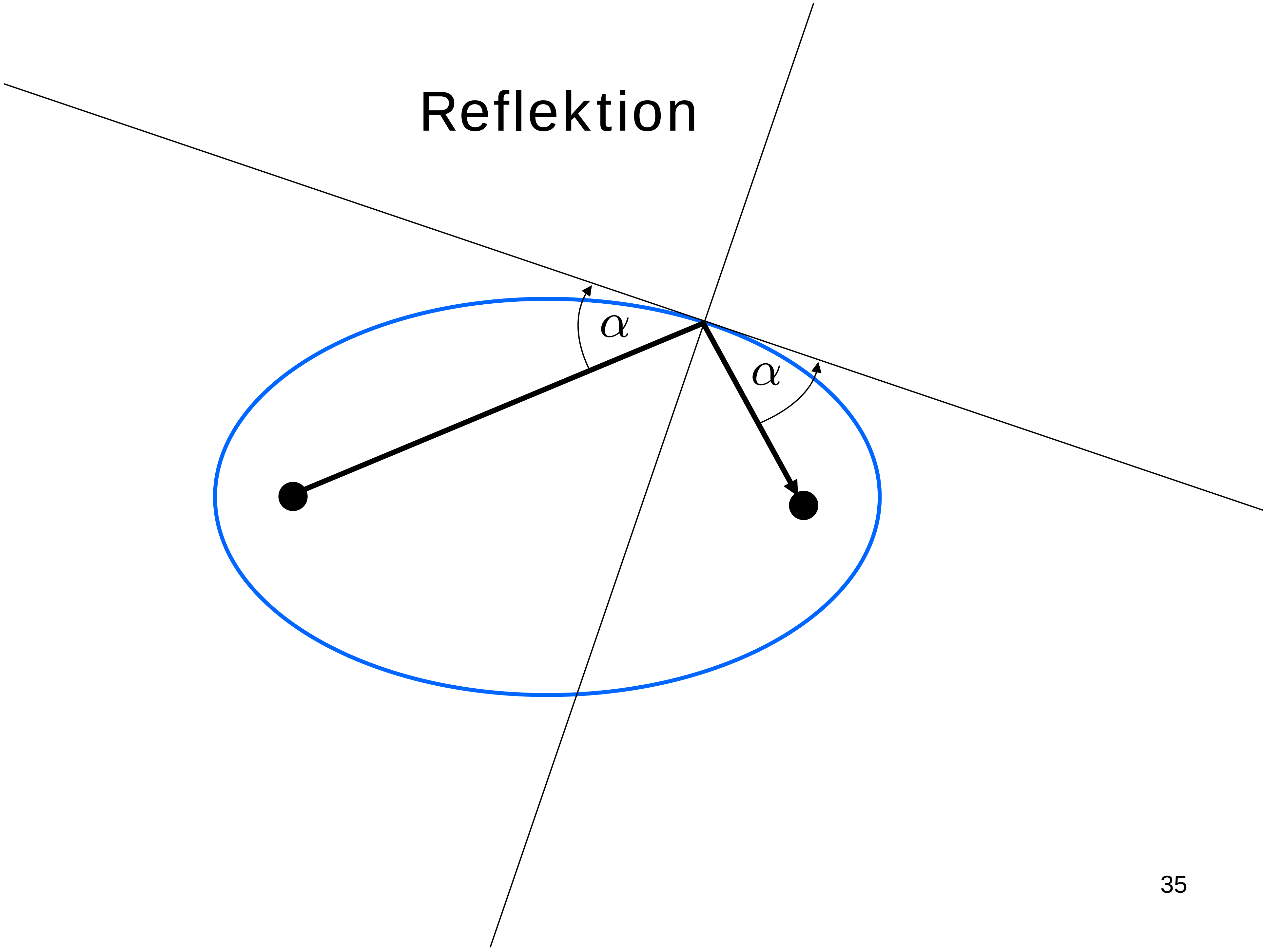
Differenz der Abstände zu 2 Punkten
ist konstant



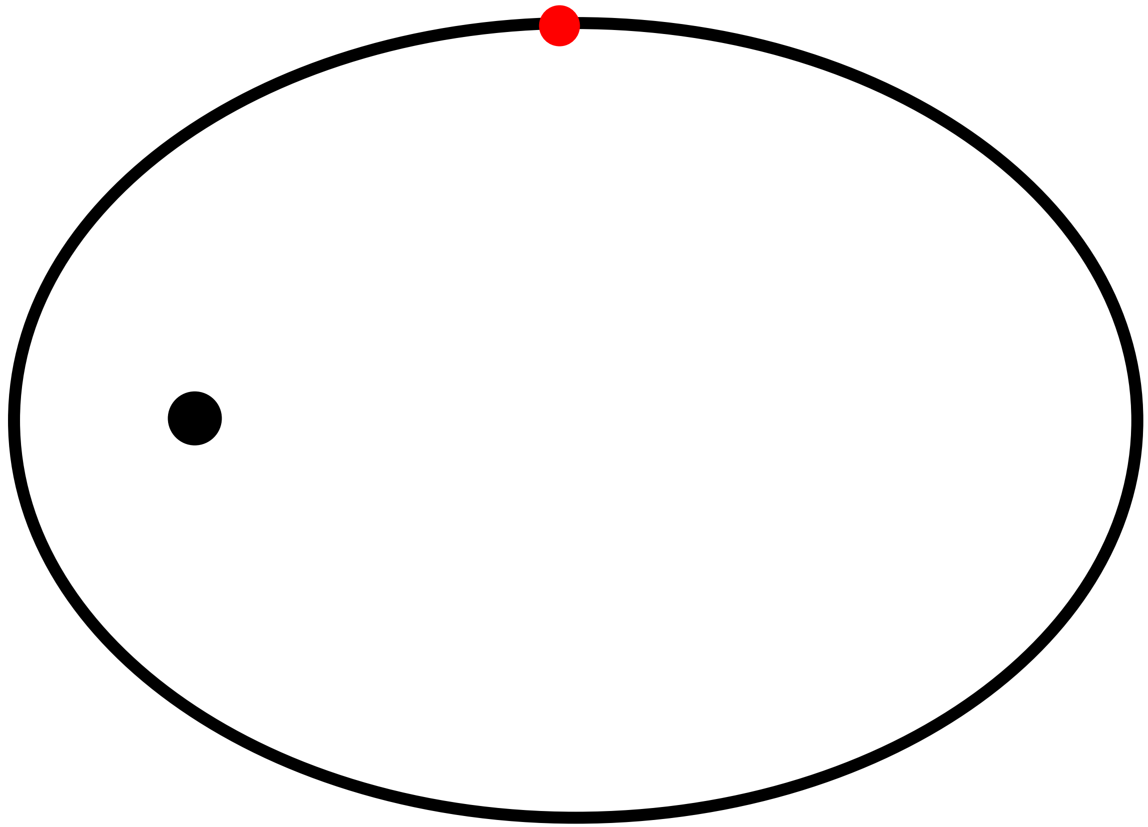


$$\frac{x^2}{a^2} - \frac{y^2}{b^2} = 1$$

Reflektion

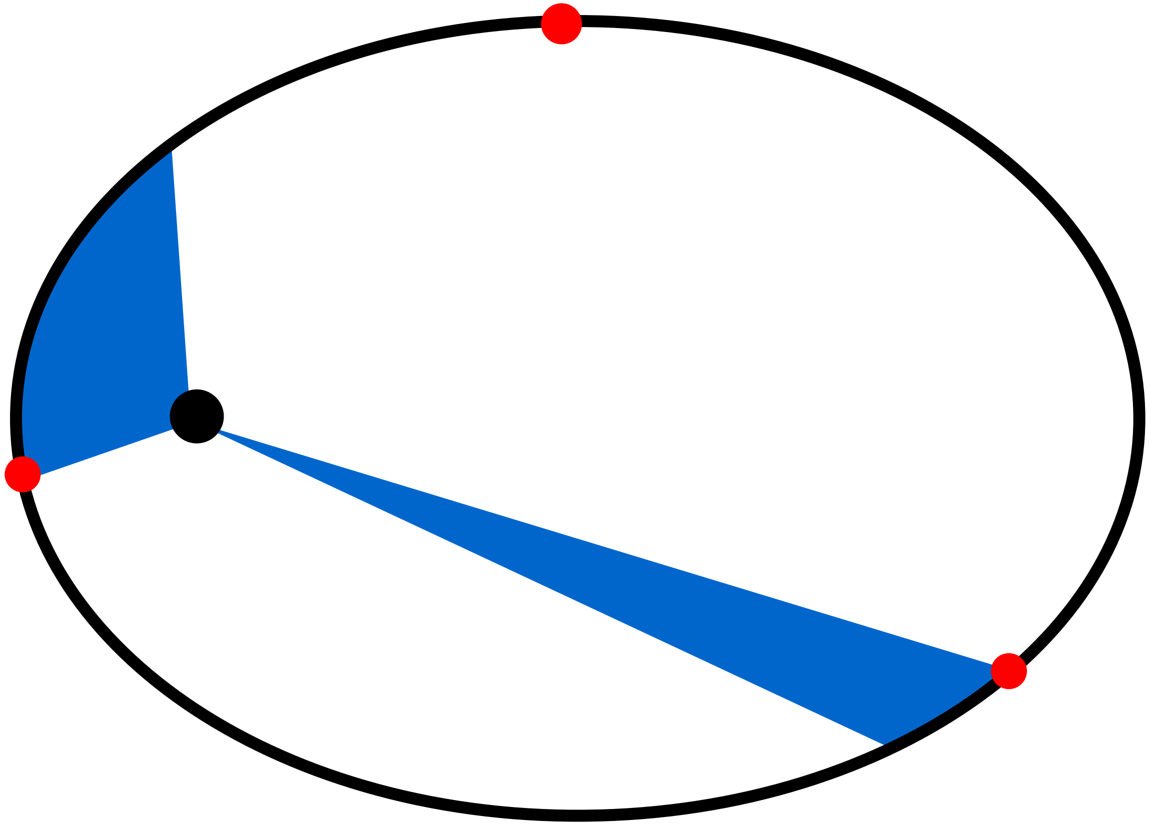


1. Keplersches Gesetz



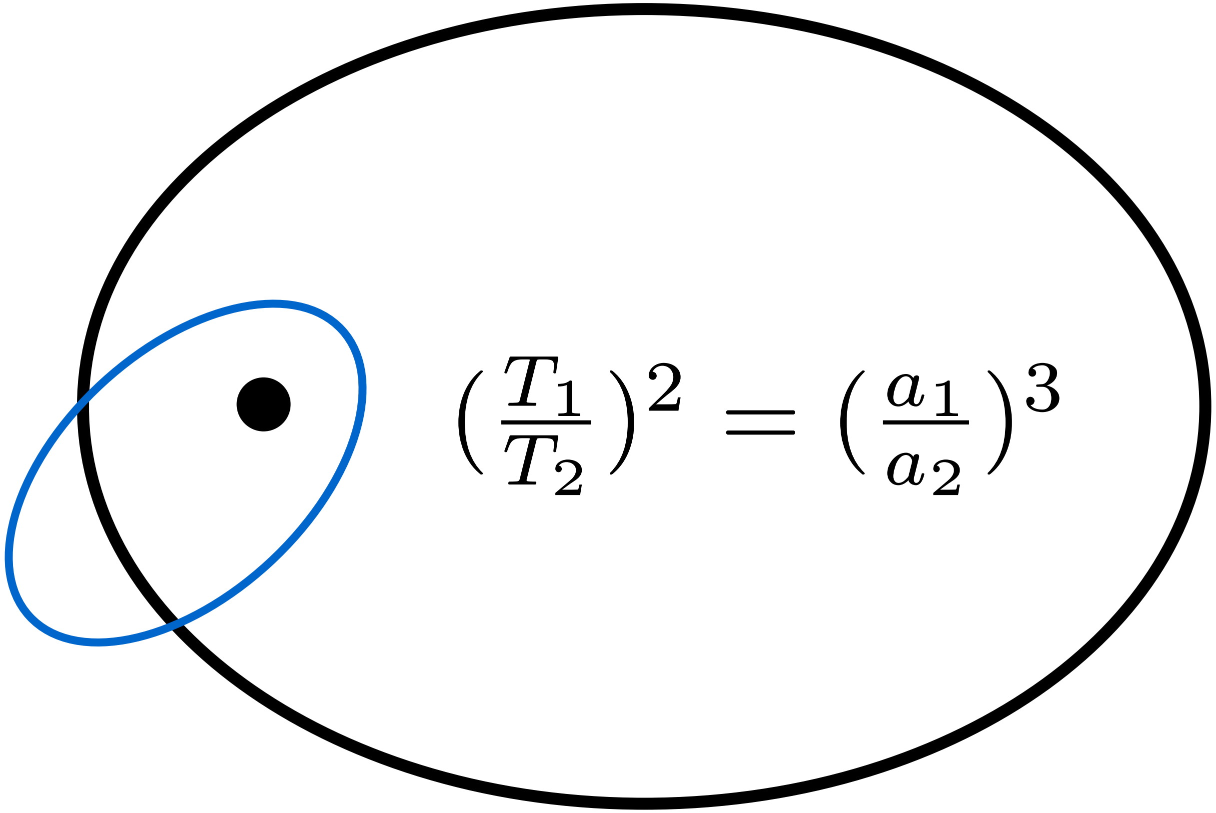
Die Planeten umkreisen die Sonne auf einer Ellipse

2. Keplersches Gesetz



In gleichen Zeiten überstreicht der Fahrstrahl gleiche Flächen

3. Keplersches Gesetz



$$\left(\frac{T_1}{T_2}\right)^2 = \left(\frac{a_1}{a_2}\right)^3$$

Die Quadrate der Umlaufzeiten verhalten sich
wie die Kuben der großen Halbachsen